

Area Seminar

Title Numerical Simulation of MHD Waves in the Solar Atmosphere

Date and Time 07/04/2014 16:00:00

Speaker Prof. Robert von Fay

Siebenburgen

Area Theoretical Physics

Venue Room No. 469

Abstract Photospheric motions, such as granular buffeting or vortices at the foot-point of magnetic flux tubes, could excite MHD waves, which propagate upwards through the solar atmosphere towards corona. In this talk I will show the results of 3D numerical simulation of generation and propagation of slow/fast magnetoacoustic and torsional Alfvén modes in the localised waveguides and determine the energy flux they carry along the magnetic field lines. Also, the new analytical approach in construction of a system of multiple magnetic flux tubes, for example, pair of open tubes and complex mixed open tube configurations with curvature and asymmetry, will be discussed. This method provides an opportunity to build a range of analytical models of magnetic field configurations that will most realistically capture magnetic structures of the lower solar atmosphere. About the speaker: Prof. Robert von Fay-Siebenburgen (Robertus) is the Director of Research at the School of Mathematics and Statistics at the University of Sheffield, UK. Prof. Robertus' group has a very high international reputation in the field of theoretical solar physics with particular expertise in the heating processes that generate and sustain the observed high temperature of the solar atmosphere by using both numerical and exact analytical methods and through observational studies using joint ground-based and satellite missions.